

1 **An investigation on the ionospheric response to the volcanic** 2 **explosion at Hunga Ha'apai in 2021**

3 Shican Qiu^{1*}, Zhanming Zhang^{1,2}, Willie Soon^{3,4}, Victor Manuel Velasco Herrera⁵,
4 Xiankang Dou^{2*}

5 ¹ Department of Geophysics, College of the Geology Engineering and Geomatics,
6 Chang'an University, Xi'an, 710054, China

7 ² School of Earth and Space Sciences, University of Science and Technology of China,
8 Hefei, 230026, China

9 ³ Center for Environmental Research and Earth Sciences (CERES), Salem, MA 01970,
10 USA

11 ⁴ Institute of Earth Physics and Space Science (ELKH EPSS), 9400, Sopron, Hungary.

12 ⁵ Instituto De Geofísica, Universidad Nacional Autónoma De México, Mexico City,
13 Mexico

14 Corresponding author: Shican Qiu (scq@ustc.edu.cn) and Xiankang Dou
15 (dou@ustc.edu.cn)

16 **Key Points:**

- 17 • **The ionospheric TEC decayed starting 10 days before the main eruption and**
18 **showed obvious fluctuations during the eruption phase.**
- 19 • **The anomaly propagated with autocorrelation-analyzed period of ~16.5 and ~8**
20 **hours during intermittent and main outbreak phases, respectively.**
- 21 • **The anomaly propagation is mainly expressed by low frequencies, with energy**
22 **concentrated in the range of 0 – 10⁻³Hz.**

23 **Abstract**

24 The Honga Ha'apai volcano eruption (20.536°S, 175.382°W in Tonga), which started
25 intermittently around December 2021 and most violently erupted on January 15, 2022, is

26 considered to be the largest volcanic outbreak in recent decades. **In this research, we**
27 **derived the ionospheric Total Electron Content (TEC) over Sanya (18.400°N,**
28 **109.600°E), Wuhan (30.530°N, 114.610°E), and Mohe (53.500°N, 122.370°E), from the**
29 **Global Navigation Satellite System (GNSS) observations.** Then we investigated the
30 coupling between the volcano eruption and ionosphere through the TEC variations. The
31 TEC anomaly decayed from 10 days before main eruption of the Honga Ha’apai volcano,
32 and showed obvious fluctuations during the eruption phase. The TEC anomaly
33 propagated periodically, with its autocorrelation-analysed period of about 16.5 hours
34 during the intermittent outbreak and about 8 hours during the main outbreak phase. Its
35 wavelet-analysed period is about 9.4 hours during the intermittent outbreak and about 9.4
36 hours and 18.8 hours during the main outbreak phase. The propagation is mainly
37 expressed in low frequencies, with energy concentrated in the range of $0 - 10^{-3}$ Hz. This
38 study highlights that the pre-eruption activities may play an important role in the coupling
39 between the volcanic eruption and ionosphere disturbances.

40 **Key words:** volcano eruption, Tonga, Honga Ha’apai, GNSS, TEC, ionosphere

41 **Plain Language Summary**

42 **In this research, we study the response of ionosphere to the volcanic explosion of Honga**
43 **Ha’apai through the Total Electron Content (TEC) over Sanya (18.400°N, 109.600°E),**
44 **Wuhan (30.530°N, 114.610°E), and Mohe (53.500°N, 122.370°E), from the Global**
45 **Navigation Satellite System (GNSS).** We investigated the coupling between the volcano
46 eruption and ionosphere through the TEC variations. We found that the TEC anomalies
47 are propagated periodically with periods of about 8–9 hours and 16–19 hours during the
48 intermittent activity phase and the main eruptive phase, respectively. We envision that
49 these signals probably reflect the nature of the Honga Ha’apai volcano eruption
50 processes.

51 **1 Introduction**

52 The Global Navigation Satellite System (GNSS) was first used for high-precision
53 global position monitoring as early as the late 1970s [*Blewitt, 1990; Dong & Bock, 1989*].
54 The Chinese Meridian Project supports GNSS observations over different latitudes
55 [*Meridian, 2022; Wang & Wei, 2007*]. The ionospheric Total Electron Content (TEC) can
56 be calculated from the GNSS ground-based observational data. The TEC is one of the
57 important physical parameters reflecting the temporal and spatial characteristics of the
58 ionosphere. It is widely used in the studies of ionospheric disturbance term correction,
59 ionospheric variation monitoring, and even long-distance emergency communication
60 [*Blagoveshchensky et al., 2005; Blewitt, 1990*]. The TEC is often affected by space
61 weather (e.g., coronal mass injection (CME), solar proton event (SPE), etc.), geological
62 activities (e.g., earthquakes, tsunamis, volcanic eruptions, etc.) and human activities (e.g.,
63 artificial explosions, large-scale radio communications, ionospheric heating, etc.)
64 [*Cahyadi et al., 2021*]. Recent studies have indeed shown and confirm that solid earth
65 activity could directly or indirectly affect the ionosphere [*Verhulst et al., 2022; Ricardo*
66 *Garza-Girón et al., 2023*].

67 Over the past 5 decades, more and more areas near earthquake zones and volcanic
68 swarms have been gradually developed as the acceleration of global urbanization and
69 population growth continues, leading to increase risks for human beings [*Unisdr, 2015*].
70 Many studies have pointed out that, in addition to obvious geological activities such as
71 volcanic eruptions and seismic fractures, geological events will be accompanied by
72 distinct electromagnetic and optical signals within the ionosphere [*Lockner, 1983;*
73 *Freund, 2000; Gokhberg & Morgounov, 1982; Leonard & Barnes, 1965*]. Therefore, the
74 coupling between geological activity and ionosphere has gradually been widely studied
75 and focused on. In 1965, a 10-minute seismic-ionospheric coupling anomaly was
76 discovered for the first time, suggesting a possible correlation between earthquakes and
77 ionospheric disturbances [*Leonard & Barnes, 1965*]. This coupling relationship is

78 subsequently determined to have occurred before the actual occurrence of geological
79 activities and events [*Gokhberg & Morgounov, 1982; Minster, 1994; Larkina, 1983*]. In
80 1982, the pre-earthquake electromagnetic radiation outburst was first observed by the
81 magnetometer on the OGO-6 satellite of the United States [*Gokhberg & Morgounov,*
82 1982]. In 1983, the very low frequency (VLF) wave was observed to be enhanced several
83 minutes before and hours after an earthquake event, through the Intercosmos-19 satellite
84 [*Larkina, 1983*]. In particular, such coupling was also found between the earthquake and
85 the TEC variations. In the study of the Northridge earthquake in 1994, it was detected
86 that the ionospheric GPS-TEC had related disturbances before and after the earthquake
87 [*Minster, 1994*].

88 Nowadays, more and more studies have confirmed that the ionosphere can be
89 effectively monitored in order to study the geological activities, through the method of
90 combining satellites-borne data with traditional ground-based observations [*Cahyadi et*
91 *al., 2022; Cussac et al., 2006; Liu, 2004; Liu et al., 2011; Satti et al., 2022; Zhang, 2008;*
92 *Zhang et al., 2009*]. Indeed, it is possible to obtain and study the variation pattern of
93 geological activity during the whole period, that is, from the precursor to the later stage,
94 based on the analyses of ionospheric parameters [*Liu, 2004; Kon et al., 2011; Liu et al.,*
95 *2006a; Liu et al., 2000; Liu et al., 2006b; Satti et al., 2022; Zhang, 2008; Zhang et al.,*
96 *2009*]. The TEC and the F2 layer critical frequency (f_0F_2) have been confirmed as
97 important indicators in the geological activity-ionosphere coupling [*Astafyeva et al.,*
98 *2011; Cahyadi et al., 2022; Liu, 2004; Liu, & Liu, 2011; Liu et al., 2011; Liu et al.,*
99 *2006a; Liu et al., 2006b; Maletkii & Astafyeva, 2022; Parrot et al., 2006a; Parrot et al.,*
100 *2006b; Pulinets, 2004; Toman et al., 2021; Zhang et al., 2009*]. In addition, more detailed
101 information, such as earthquake location, intensity, and local tectonic orientation, can be
102 determined from the propagation pattern of the ionospheric disturbances [*Cahyadi et al.,*
103 *2021; Cahyadi et al., 2022; Afraimovich, 2001; Le et al., 2011; Liu et al., 2006a; Liu et*
104 *al., 2000*]. The mechanism for this coupling has been analyzed, and candidates such as

105 the ionizing radiation model and conductivity model have been proposed [*Lockner*, 1983;
106 *Gokhberg & Morgounov*, 1985; *Parrot et al.*, 2006a; *Parrot et al.*, 2006b; *Pulinets*, 2004;
107 *Zhang*, 2008]. Recently, *Velasco Herrera et al.* (2022) has reviewed the earthquake–
108 ionosphere relationship as potential precursors for forecasting major strong earthquakes
109 over major fault zones/regions of the world.

110 As for the specific direction of volcano–ionosphere coupling research, it has been
111 established that there is a correlation between ionospheric and volcanic activity, with
112 TEC variation as an effective diagnostic parameter [*Aoyama et al.*, 2016; *Cahyadi et al.*,
113 2021; *Cahyadi et al.*, 2020; *Heki & Fujimoto*, 2022; *Li et al.*, 2016; *Lin*, 2017; *Liu et al.*,
114 2017; *Maletckii & Astafyeva*, 2022; *Manta et al.*, 2021; *Pandara et al.*, 2021; *Saito*, 2022;
115 *Shults et al.*, 2016; *Toman et al.*, 2021; *Verhulst et al.*, 2022; *Zhang et al.*, 2022].

116 **Meanwhile, observation from multiple independent instruments can show more details**
117 **and bring a fuller picture about the nature of the coupling** [*Verhulst et al.*, 2022]. Thus, a
118 combination observational data from both ground–based and space–based method is
119 particularly important [*Heki & Fujimoto*, 2022; *Matoza et al.*, 2022; *Wright et al.*, 2022].
120 The combined analyses had found abnormal propagation that has multiple wave
121 characteristics such as acoustic wave, infrasonic wave, ultrasonic wave and gravity wave,
122 and periodic harmonic oscillation exists in volcanic eruptions [*Amores et al.*, 2022;
123 *Aoyama et al.*, 2016; *Cahyadi et al.*, 2021; *Heki & Fujimoto*, 2022; *Lin*, 2017; *Lin et al.*,
124 2022; *Liu et al.*, 2022; *Liu et al.*, 2017; *Manta et al.*, 2021; *Nakashima et al.*, 2016;
125 *Pandara et al.*, 2021; *Maletckii & Astafyeva*, 2022; *Matoza et al.*, 2022; *Ricardo Garza-*
126 *Girón*, 2023; *Shults et al.*, 2016; *Kubota et al.*, 2022; *Toman et al.*, 2021; *Verhulst et al.*,
127 2022; *Watson et al.*, 2022; *Wright et al.*, 2022; *Zhang et al.*, 2022]. By analyzing the
128 spectrum of ionospheric anomalies, different wave propagation characteristics and
129 volcanic eruption types can be distinguished [*Heki & Fujimoto*, 2022; *Li et al.*, 2016; *Liu*
130 *et al.*, 2017; *Nakashima et al.*, 2016]. In addition, the intensity of ionospheric anomalies
131 has also been confirmed to be significantly correlated with volcanic eruption intensity

132 and plume height [*Cahyadi et al.*, 2020; *Manta et al.*, nbo; *Shults et al.*, 2016; *Toman et*
133 *al.*, 2021]. Based on the study of anomalous propagation characteristics in the ionosphere,
134 a variety of information can be obtained, including the location of volcanic source, the
135 scale of volcanic eruption, the height of plume, the rate of material ejection and the
136 quality of ejection materials [*Cahyadi et al.*, 2020; *Li et al.*, 2016; *Maletckii & Astafyeva,*
137 *2022*; *Shults et al.*, 2016; *Watson et al.*, 2022]. In the study of volcanic eruption activity
138 cycles, it has been found that ionospheric anomalies can occur several days before the
139 main eruption event [*Li et al.*, 2016; *Pandara et al.*, 2021; *Toman et al.*, 2021].
140 Particularly, the enhancement of TEC has been observed to occur during the main phase
141 of volcanic eruption [*Toman et al.*, 2021].

142 Overall, the specific physical mechanism of volcanic–ionosphere coupling is still
143 highly uncertain, which is why a large number of observational evidence is both
144 necessary and important [*Lockner*, 1983; *Gokhberg & Morgounov*, 1985; *Parrot et al.*,
145 *2006a*; *Parrot et al.*, 2006b; *Pulinets*, 2004; *Zhang*, 2008]. For our case study, we
146 consider the Hunga Ha’apai volcanic eruption in Tonga (20.536°S, 175.382°W), which is
147 one of the strongest volcanic eruptions in recent years with a start date on December 20,
148 2021 universal time (UT) [*Poli & Shapiro*, 2022; *Matoza et al.*, 2022]. After the first
149 eruption, volcanic activities continued to come and go but weaken for about two weeks
150 [*INGV*, 2022; *Kusky*, 2022; *NPR*, 2022]. After that, the eruption resumed on January 13,
151 2022, and the largest outbreak was observed at around 4:00 UT on January 15 when the
152 top umbra cloud reached a maximum diameter of 500 km [*GVP*, 2022; *INGV*, 2022;
153 *Kusky*, 2022; *NASA*, 2022; *NPR*, 2022]. Current analysis suggests that it has a Volcanic
154 Explosivity Index (VEI) of 5 or 6 or even higher [*INGV*, 2022; *NASA*, 2022; *Poli &*
155 *Shapiro*, 2022]. Proud et al. (2022) recently confirmed that the volcanic cloud for this
156 event reached the extreme height of 57 km at its highest extent.

157 The main purpose of this study is to investigate the coupling between volcanic
158 eruption and the ionosphere. The vertical TEC is used as the indicator of ionospheric

159 disturbance, through the ground-based GNSS observations over Sanya, Wuhan, and
 160 Mohe stations, of the Chinese Meridian Project. The Tonga volcanic eruption is divided
 161 into two stages: the first stage is mostly with intermittent eruption activities, lasting about
 162 two weeks since December 20, 2021 and then weakened around January 3; the second
 163 stage is the main outbreak phase, starting on January 13 with a large-scale concentrated
 164 eruption persisting for many consecutive days. This study can offer as a supplementary
 165 method to tackle certain cases of eruptions when there are significant data gaps or
 166 simply with missing data.

167 **2 Data and Method**

168 In this study, the ground-based GNSS observational data, the Broadcast Ephemeris
 169 Products, the Differential Code Biases data (DCB) between the satellite and ground
 170 station, and the volcanic eruption development data are processed.

171 Based on the GNSS data, the TEC can be estimated using the ionospheric delay
 172 term. According to the theory, the propagation velocity of the ranging code in the
 173 ionosphere is the group velocity V_g , from which the geometric distance ρ between the
 174 satellite and the ground receiver satisfies the following equation:

$$175 \quad \rho = \int_{\Delta t} V_g dt = \int_{\Delta t} (c - 40.28c \frac{N_e}{f^2}) dt = c\Delta t - \frac{40.28}{f^2} \int_{\Delta t} cN_e dt, \quad (1)$$

176 where Δt is the propagation time of satellite signal passing to the ground receiver [Feng,
 177 2020].

178 The signal path through the ionosphere can be approximated by a cylinder-shaped
 179 volumetric tube, the TEC expression can be obtained from the definition as:

$$180 \quad TEC = \int N_e ds, \quad (2)$$

181 Then the delay magnitude for ionospheric effects on phase (I_p) and ranging code
 182 (I_g) can be expressed as follows:

$$183 \quad I_p = -I_g = \frac{40.28}{f^2} TEC, \quad (3)$$

184 In the dual-frequency GNSS system, there will be clock offsets, multi-path delays,
 185 hardware delays, and observational noises. Therefore, the following equations are used
 186 for the calculation:

$$187 \quad TEC_P = \alpha P = \alpha(P_1 - P_2) = TEC_{abs} + DCB, \quad (4)$$

$$188 \quad TEC_L = \alpha L = \alpha(L_1 - L_2) = TEC_{abs} + DCB + \frac{1}{\alpha}(\lambda_1 N_1 - \lambda_2 N_2), \quad (5)$$

$$189 \quad L_1 = \lambda_1 \phi_1, \quad (6)$$

$$190 \quad L_2 = \lambda_2 \phi_2, \quad (7)$$

191 where

192 TEC_P ~ differential pseudorange TEC, also known as absolute TEC;

193 P_1 and P_2 ~ pseudorange of the dual-frequency observations;

194 TEC_{abs} ~ true value of TEC;

195 DCB ~ differential code deviation;

196 TEC_L ~ carrier phase TEC, also known as relative TEC;

197 λ_1 and λ_2 ~ the wavelengths of phase carriers L_1 and L_2 , respectively;

198 N_1 and N_2 ~ the whole-period ambiguities of phase carriers L_1 and L_2 ,
 199 respectively;

200 and

201 ϕ_1 and ϕ_2 ~ the carrier phase [Feng, 2020].

202 Base on the calculations of TEC, a comparison of TEC for each day with the median
 203 value for that period can be obtained. Then Fourier and autocorrelation analyses are
 204 carried out for examining the abnormal variation of TEC. The Fast Fourier Transform
 205 (FFT) is performed as

206

$$207 \quad X_m = \sum_{n=0}^{N-1} x_n e^{-inm\frac{2\pi}{N}} \quad (0 \leq m \leq N - 1), \quad (8)$$

$$208 \quad \begin{cases} X_l = G_l + W_N^l H_l \\ X_{\frac{N}{2}+l} = G_l - W_N^l H_l \end{cases} \quad \left(0 \leq l \leq \frac{N}{2} - 1\right), \quad (9)$$

209 where

210 $X_m \sim$ the finite discrete spectrum;

211 $x_n \sim$ the finite discrete signal;

212 G_l and $H_l \sim$ the discrete spectra of the even and odd terms, respectively;

213 and

$$214 \quad W_N^l = e^{-i\frac{2\pi}{N}} \quad [Cheng, 2010].$$

215 When the calculation is simplified to only one term, its finite discrete spectrum is the
 216 finite discrete signal itself, and the inverse derivation can be performed. The
 217 autocorrelation function is calculated using the following equation

$$218 \quad r_{xx}(\tau) = \sum_{n=-\infty}^{\infty} x_n x_{n-\tau} = x_n * x_{-n}, \quad (10)$$

219 where $r_{xx}(\tau)$ is the autocorrelation function with respect to the time shift τ and x_n is
 220 the calculated signal [Cheng, 2010].

221 As shown in Eq. (10), the signal amplitude will gradually deviate from the
 222 corresponding value with the time shift, resulting in the reduction of the amplitude of the
 223 correlation function. However, when the signal has quasi-periodicity, it will reveal an

224 approximately equally spaced amplitude. The amplitude attenuation from the correlation
225 results will be weakened, and there will appear obvious sub-peaks that are slightly
226 smaller than the main peak. The number of subpeaks is related to the number of original
227 peaks in the calculated signal, and can be inferred from the correlation function.

228

229 **3 Ionospheric Response to Volcanic Eruption**

230 In this study, we speculate or assume that the pre-eruption phase will cause dense
231 earthquakes and release significant amounts of gas from magma capsule as the Figure 1 below
232 sketched. These processes will change the composition over local atmosphere and generate both
233 gravity and seismic waves, and hence will further affect the ionosphere [*Sigurdsson et al.*,
234 2015]. We believe that there are a variety of gases dissolved in magma capsule like carbon
235 dioxide (CO_2), sulfur dioxide (SO_2), hydrogen sulfide (H_2S), carbon monoxide (CO), hydrogen
236 chloride (HCl) and so on [*Sigurdsson et al.*, 2015]. There are sufficient evidence to suggest that
237 these gases which leak from underground fissure during the pre-eruption period can change the
238 temperature of upper atmosphere and lower the electron density [*Cnossen, 2022; Qian et al.*,
239 2013; *Roble & Dickinson, 1989*]. In the meantime, during the peak phase of volcanic eruptions
240 much more gravity wave, seismic wave, acoustic wave and infrasonic wave will be generated
241 [*Ricardo Garza-Girón, 2023*]. Although this effect is not global and have low strength, the local
242 variation will propagate outward by the dynamic action of the upper atmosphere.

243 The data for the first stage (from December 10 to 31, 2021) are shown in Figure 2(a–l). The
244 horizontal coordinate is Day of year (Doy) and the ordinate is Total Electron Content Unit
245 (TECU), where 1 TECU means 10^{16} electrons per unit area. It can be observed that at the
246 beginning of the December 20 (Doy354), the fitted values deviate obviously and significantly
247 from the median. The TEC fluctuates significantly, and the largest peak value of TEC reaches 50
248 TECU, much higher than the average daily maximum of 20–30 TECU. Figure 2(m–t) exhibits
249 the observed results from January 10 to 25, 2022. The data shows a disturbance on January 15,
250 2022 (Doy15), but the amplitude is distinctly smaller than that in Figure 2(a–l).

251 The variations of the anomaly values obtained by subtracting the observed TEC from the
252 median are shown in Figure 3. Similar characteristics of the TEC anomalies occur over different
253 stations. In the period before the formal eruptions for both stages, the TEC decays in different
254 degrees. The attenuation amplitude in the intermittent stage is much larger, with obvious
255 characteristics of oscillation and fluctuation after the outburst starts. Meanwhile, there are
256 obvious differences between stations. According to the Space Environment Prediction Center
257 (SEPC), there is a moderate geomagnetic storm occurred at January 14 and 15, which may
258 induce ionospheric responses [*SEPC, 2022a, 2022b*]. The effect of a storm is wide and global

259 with minimal attenuation between different stations. On the contrary, the anomalies caused by a
 260 volcanic eruption will decay with the propagation obviously, and have a distinct variation from
 261 quiet period and clear fluctuation characters. From Figure 3, we can find out that the intensity
 262 of anomaly decreases with the increase of distance to Honga Ha'apai. This is similar to
 263 previous study. According to Verhulst et al. (2022), they found that the TIDs during this period
 264 are not of auroral origin, i.e., not related to the geomagnetic storm [Verhulst et al., 2022].
 265 Therefore, we believe that the impact of the Honga Ha'apai eruption is much more prominent
 266 than that from the moderate magnetic storm.

267 Therefore, the TEC anomaly over Sanya station, which is closer to Tonga, is more intense
 268 compared to that of Wuhan and Mohe stations. But there are anomalous peaks in Wuhan and
 269 Mohe, which do not appear in Sanya. On December 20 (Doy354) and 21 (Doy355), distinct TEC
 270 enhancement occurs over Wuhan and Mohe, while the TEC oscillates obviously over Sanya on
 271 that time. On January 15 (Doy15) and 16 (Doy16), the fluctuation of TEC anomaly shows
 272 similar distinct waveform characteristics. If we consider the two maxima between Doy 15 and
 273 Doy 16 (pointed out by the blue asterisks in Figure 3b, 3d, and 3f), the calculated time difference
 274 is about 21.7 hours for Sanya, 19.9 hours for Wuhan, and 20.5 hours for Mohe, respectively.
 275 These two fluctuations may be generated by the same disturbance originated from the eruption,
 276 but they arrive at observed stations from two opposite directions over the globe. Based on the
 277 coordinates of Sanya station in Hainan (18.400°N, 109.600°E), Wuhan in Hubei (30.530°N,
 278 114.610°E), Mohe in Heilongjiang (53.500°N, 122.370°E), and Honga Ha'apai volcano in Tonga
 279 (20.536°S, 175.382°W), the time difference of the anomaly arriving at observed stations from two
 280 directions can be evaluated using the ellipsoid model of the Earth.

281 The Earth's ellipsoidal model is given as follows:

$$282 \quad r = \sqrt{(x_2 - x_1)^2 + (y_2 - y_1)^2 + (z_2 - z_1)^2}, \quad (11)$$

$$283 \quad S = r + \frac{r^3}{24*k^2} + \frac{3*r^5}{640*k^4}, \quad (12)$$

284 and the time difference can be calculated by

$$285 \quad \Delta T = T_1 - T_2 = \frac{H - S}{v} - \frac{S}{v}, \quad (13)$$

286 where

287 (x_1, y_1, z_1) and $(x_2, y_2, z_2) \sim$ the coordinate positions of the two points in the Cartesian
 288 coordinate system;

289 $r \sim$ the straight–line distance between the two points;

290 $S \sim$ the distance between the two points on the ellipsoid surface;

291 $k \sim$ the curvature radius of the arc between the two points;

292 T_1 and $T_2 \sim$ the time for the propagation from the two opposite directions;

293 $v \sim$ the anomaly propagation velocity;

294 and H is the circumference of the Earth (e.g., $H = 2\pi R$, and R is the Earth radius).

295 Setting the Earth’s eccentricity of about 0.082, the semi–major axis of the Earth’s ellipsoid
 296 of 6378.137 km, the ionosphere height of 350 km, and v ranges 300~350 *km/s* according to
 297 the recent reference, the calculated average time differences is $\Delta T_{Sanya} = 19.58 \pm 3.36$ hours,
 298 $\Delta T_{Mohe} = 17.97 \pm 3.36$ hours and $\Delta T_{Wuhan} = 19.35 \pm 3.36$ hours, where the measured time
 299 differences is nearly within the calculated range.

300 **4 Discussion**

301 As seen in Figure 3, the fluctuation characteristics of the anomalies are consistent among
 302 different stations, but a certain degree of attenuation may occur at some stations. For analysis of
 303 the characteristics of the volcanic eruption signals, we make further processing on the anomalous
 304 signals to explore the coupling between the volcanic activities and ionospheric variations.

305 **4.1 The Autocorrelation Analysis and Wavelet Analysis**

306 The autocorrelation has little effect on the periodicity of the original signal, so the results
 307 can be used to analyze the periodicity of the anomalies in the first or initial stage of analysis. The
 308 autocorrelation of the TEC anomaly over Sanya during the intermittent stage is shown in Figure
 309 4a, which reveals that the maximum peak is not prominent when compared to other smaller
 310 peaks. This result indicates that with time shift, the differences in TEC anomalies caused by each
 311 eruption are similar throughout stage 1. However, the interval of each volcanic eruption is not
 312 completely consistent, so in the process of calculating autocorrelation and continuous time shift,
 313 there will be a smaller peak similar to that with large correlation. Since the interval of each

314 eruption is not completely consistent, smaller peaks similar to the larger correlations will occur
315 in the process of calculating autocorrelation with continuous time shift.

316 As the time shift increases, it can be seen that the image is clearly divided into two regions,
317 and the amplitude corresponding to the first 5000 points in the first half of the autocorrelation
318 function is small overall. This is because the anomalous features in the precursor phase may
319 differ significantly from those in the outbreak phase, and the correlation becomes worse as the
320 time shift increases. And according to the principle of autocorrelation, there are about 4 or 5 sub-
321 peaks in the correlation function results that are close to the main frequency peak. This indicates
322 that there are also a number of correlated peaks in the original calculated signals, and these peaks
323 are the regions where the TEC anomaly changes the most, so it can be tentatively judged that
324 there were at least 3–4 outbursts of similar size to the December 20th's intermittent outburst
325 stage.

326 In addition, as shown in Figure 4b, during the main eruption phase, the detected main peaks
327 are 4–10 times larger than the neighboring smaller peaks over Sanya station. This result indicates
328 that the eruption size in this stage is the largest and most concentrated during the violent eruption
329 phase. Periodic analysis of the portion of the autocorrelation function with peaks greater than
330 10^4 and suggests the existence of a 16.5-hour periodic variation characteristic for stage 1 and an
331 8-hour periodic variation characteristic for stage 2.

332 Furthermore, Figure 4c (for Wuhan) and 4e (for Mohe) show similar results during the
333 intermittent stage to Figure 4a. And in Figure 4d and 4f, we can find a similar pattern of
334 autocorrelation results during the main eruption phase as in Figure 4b. However, the amplitude
335 of the autocorrelation results for the Mohe and Wuhan become significantly reduced, with about
336 one order of magnitude smaller than that of Sanya. This discrepancy indicates that the signal
337 intensity is attenuated during the propagation.

338 On the other hand, the wavelet transform can perform high and low filtering and time-
339 frequent domain analysis with higher resolution [*Chen et al.*, 2021; *Hubbard*, 1998; *Pan et al.*,
340 2008]. Therefore, more detailed information on the signals can be obtained by adopting wavelet
341 analysis. With this property, the signal period can be extracted while at the same time the signal
342 can be filtered according to the selected wavelet and scale. The result of wavelet analysis of the
343 TEC anomaly during the intermittent stage is shown in Figure 5a. We can find that there is a

344 periodic oscillation of 9.4 hours in the intermittent stage. In the results of the main eruption
345 stage, we found the periodic variations of 9.4 hours and 18.8 hours (as shown in Figure 5b). The
346 results of wavelet analysis shown in Figure 5 are basically consistent with those of
347 autocorrelation analysis from Figure 4a and 4b.

348 4.2 The Fourier Analysis

349 From the wavelet transform results with the high frequency part, it can be observed that
350 signals seldom concentrate in the high frequency regime (Figure 6a and 6b). Considering the
351 volcanic activity is continuous in the selected time period, the low frequency part of the signal
352 can be further analyzed by the Fourier analysis. Compared with wavelet transform, the Fourier
353 transform is more suitable to deal with the whole signal. Based on these characteristics, the
354 Fourier transform can be used to highlight frequency ranges where the signal energy is more
355 concentrated [Chen *et al.*, 2021; Iyer, 1968; Toman, 1966; Troyan & Kiselev, 2010]. The Fourier
356 analysis for the anomaly of TEC during the intermittent stage and the main eruption stage is
357 performed, with the output amplitude spectra shown in Figure 6c and 6d. It can be confirmed
358 from Figure 6c that the energy of the anomalous fluctuations during the intermittent stage is
359 mainly concentrated in the frequency range of an order of $0 - 10^{-3}$ Hz. Figure 6d shows the
360 amplitude spectrum of the TEC anomalies during the main phase. The energy of the anomalous
361 fluctuations during the main stage is also concentrated in the frequency range of an order of $0 -$
362 10^{-3} Hz. These results indicate that the anomalous fluctuations caused by the eruption are
363 mainly low-frequency waves. However, the maximum amplitude of the wave of the main phase
364 is smaller than that of the intermittent stage. Therefore, it is probably that some physicochemical
365 processes occurring during the intermittent phase that have a greater effect on the energy of the
366 fluctuations. The precursor in the early stage of the eruption has a greater impact on ionospheric
367 disturbances, and the difference in precursory activities between stage 1 and stage 2 would
368 possibly lead to the distinct fluctuations during the two stages.

369 4.3 Possible Coupling between the Volcanic Eruptions and TEC Variations

370 The analysis here reveals that the physicochemical processes at different stages of volcanic
371 activity are distinct, indicating different mechanisms for their effects on the TEC variations.
372 Figure 3 shows that in the precursory stage, the TEC exhibits a significant attenuation. For the
373 Surtseyan and Plinian eruptions, during the precursor stage (as showing in Figure 1), large-scale

374 escape of gases from the magma storage occurs, such deeply-rooted gases are released by crustal
375 rupture, and underground materials sublimate fast. Therefore, the magma does not overflow at
376 this precursor stage yet. During the main eruption stage, it is often accompanied by the eruption
377 of lava, volcanic ashes, and pyroclastic flows. The TEC also exhibits different abnormal
378 characteristics from the previous stage. The variations of TEC first show distinct decline and
379 then there are enhancement and attenuation alternately. And the enhancement propagates with
380 obvious periodicity.

381 In addition, in the comparative analysis of intermittent stage and main eruptions stage, the
382 variation of the peak value of TEC anomaly is revealed. In the intermittent eruption stage, the
383 maximum TEC anomaly can reach 50 TECU, and the magnitude of TEC appears to be elevated
384 throughout this period. During the main eruption phase, the maximum TEC anomaly reaches 30–
385 40 TECU, and the TEC, however, does not exhibit any positive correlation between its increased
386 magnitude and the eruption intensity. These results can match our hypothesis and observation
387 that the greater effect of early volcanic activity on TEC is mainly because of the escaping
388 materials/gasses during the precursor period. The lower density and the faster emission rate of
389 the early escape materials, a large amount of gasses has already been depleted during the
390 intermittent eruption phase. Since there is already reduced amount of residual matter in the
391 magma storage, it will probably cause a smaller effect on TEC during the main eruption phase.

392 **5 Conclusion**

393 From the analysis carried out in this paper, it can be confirmed that the high sensitivity of
394 TEC in relation to different stages of volcanic eruption is of great significance for studying the
395 mechanism of the coupling between volcanic activity and ionosphere. Based on the results from
396 this study, the following conclusions can be drawn:

397 1. From 10 days before the main eruption, obvious attenuation appeared in the TEC
398 profiles. During the main phase, the TEC anomaly alternated between attenuation and
399 enhancement. And the anomaly showed fluctuations in both stages.

400 2. The anomaly propagated periodically, with its autocorrelation-analysed period of about
401 16.5 hours during the intermittent outbreak and about 8 hours during the main outbreak phase. Its
402 corresponding wavelet-analysed period is about 9.4 hours during the intermittent outbreak and
403 about 9.4 hours and 18.8 hours during the main outbreak phase.

404 3. The anomaly propagation is mainly dominated and expressed by low frequencies, with
405 energy concentrated in the range of $0 - 10^{-3}$ Hz.

406 4. The anomaly generated by the Honga Ha'apai eruption arrived successively at observed
407 stations from two opposite directions, with average differences of about 19.58 hours (Sanya),
408 19.35 hours (Wuhan), 17.97 hours (Mohe). Based on actual GNSS-TEC observation, the
409 differences during this eruption are about 21.7 hours (Sanya), 19.9 hours (Wuhan), 20.5 hours
410 (Mohe).

411 5. During the precursor stage of the eruption, there were large imprints on the TEC, while
412 the amplitude of TEC anomaly during the main eruption phase is much smaller. This indicates
413 that activities in pre-eruption stage like gas releasing have a greater impact on TEC than activities
414 such as explosion generated by the eruption. This discrepancy may be related to the physical and
415 chemical processes generated by a large number of volatile substances before the main phase of
416 the eruption. This result can provide more evidence for the study of the coupling process
417 between volcanic activity and ionosphere.

418 6. According to the results of our autocorrelation analysis, at least another five to six
419 eruptions might have occurred with similar size to the case on December 20 during the
420 intermittent phase.

421 **Open Research**

422 The ground-based GNSS observations are available from the Space Environment Ground-based
423 Integrated Monitoring Network of Chinese Meridian Project database
424 (<http://meridianproject.ac.cn/>). The GNSS Broadcast Ephemeris Products are downloaded from
425 the IGS Data Center of Wuhan University (<http://www.igs.gnsswhu.cn/>). The DCB data of
426 GNSS observation system come from NASA's Crustal Dynamics Data Information System
427 website (<https://cddis.nasa.gov/index.html>). The volcanic eruption data are mainly from the
428 Smithsonian Institution's Global Register of Volcanic Activity (<https://volcano.si.edu/>). The
429 Geomagnetic activity information come from the Space Environment Prediction Center
430 (<http://www.sepc.ac.cn/>). The seismic information are mainly from the U.S. Geological Survey
431 (<https://earthquake.usgs.gov/>).

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438 **for the Geomagnetic activity information and the U.S. Geological Survey for the seismic**
439 **information.**

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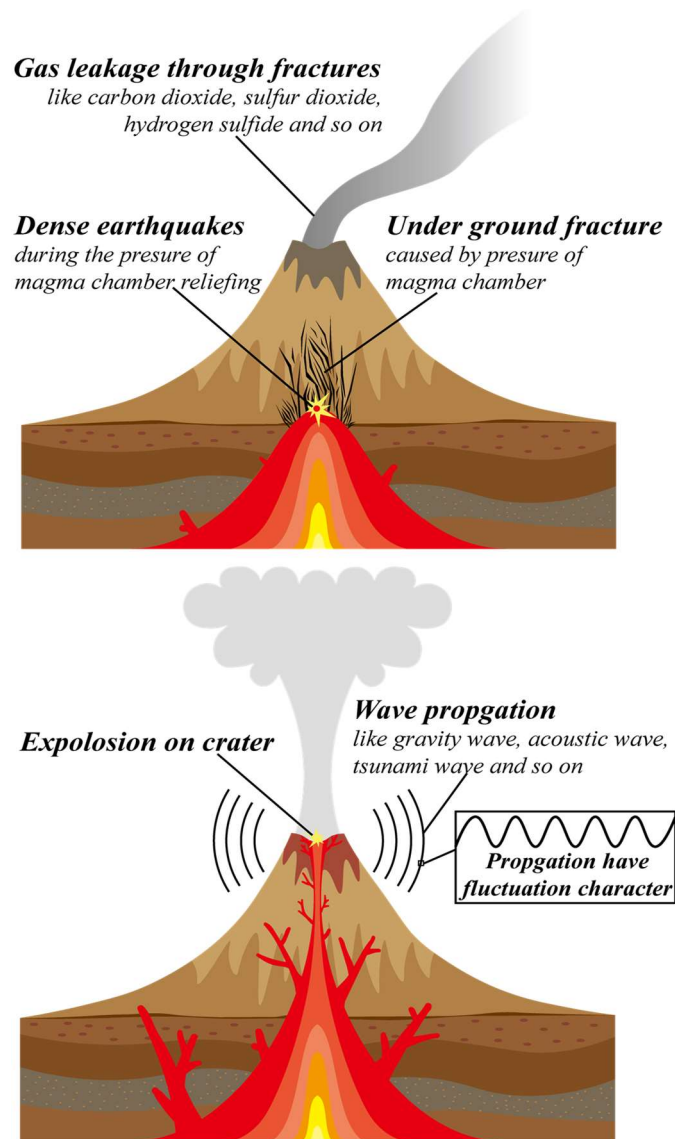
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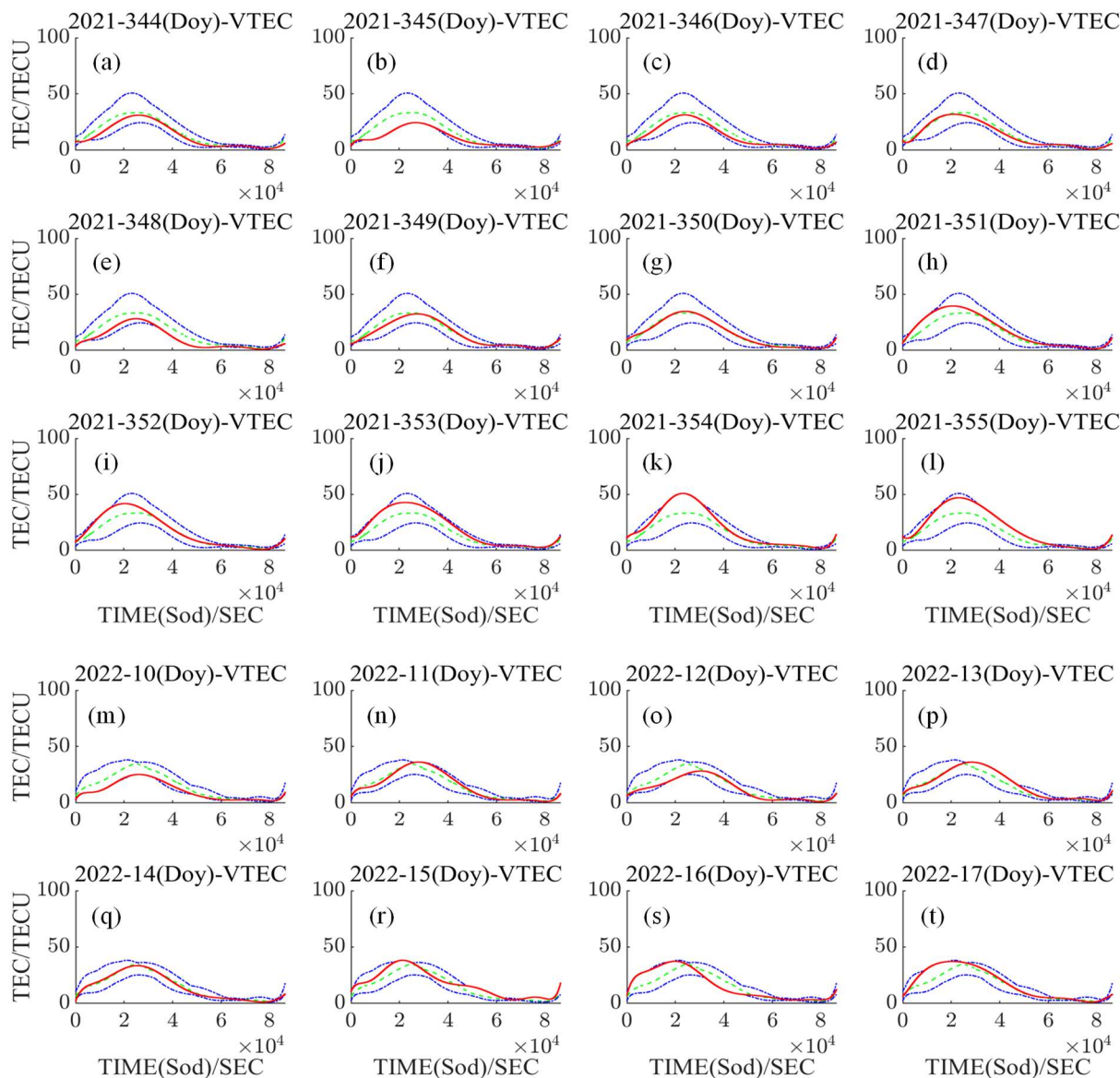
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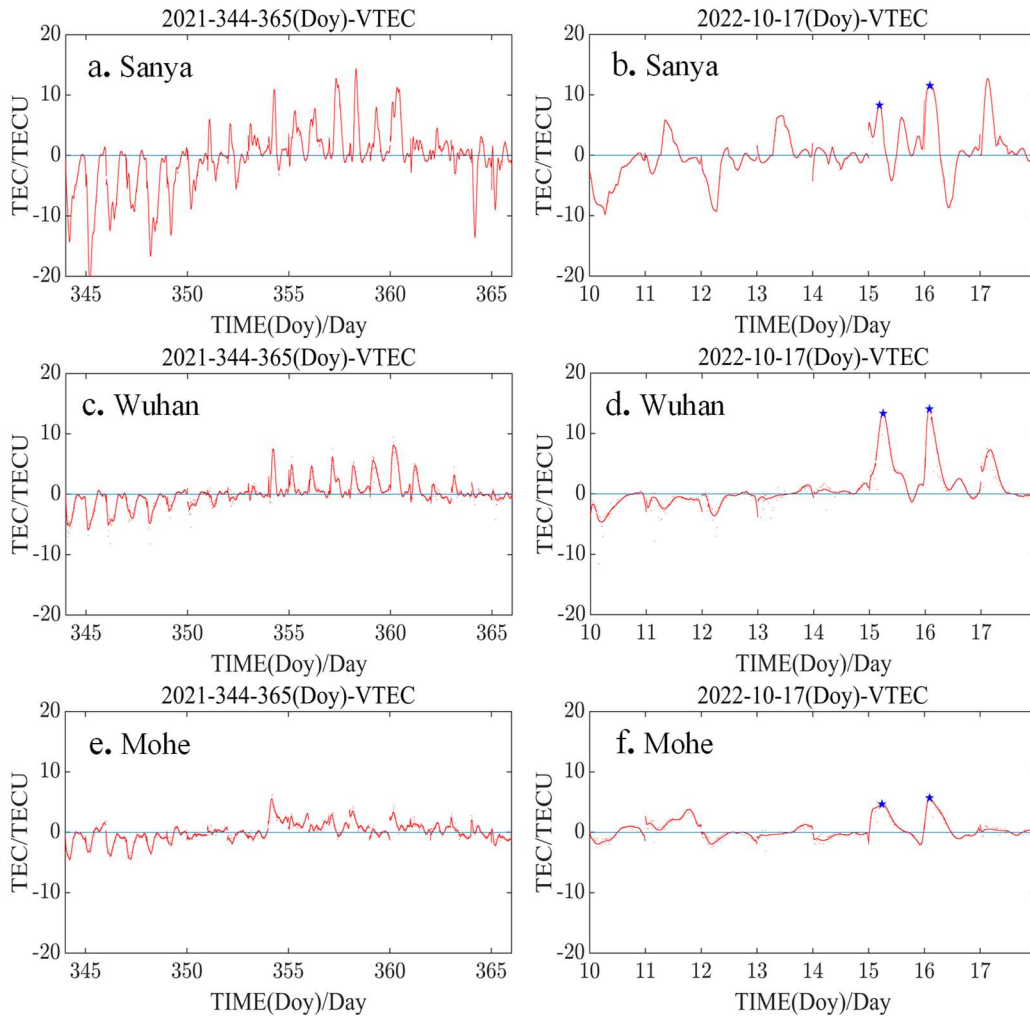
633 **Figure 1.** Schematic diagram of volcanic-related processes during the pre-eruption and eruption phases. The upper
 634 (top) figure is for the pre-eruption phase, and the lower (bottom) figure is for the eruption phase. In this study, we
 635 assume that the pre-eruption phase will trigger dense earthquakes and cause the release of significant amounts of gas
 636 from magma capsule as the sketch is indicating. These processes will change the composition over local atmosphere
 637 and generate both gravity wave and seismic wave, and they will further affect the ionosphere [Sigurdsson et al.,
 638 2015]. We believe that there are a variety of gases dissolved in magma capsule like carbon dioxide (CO_2), sulfur
 639 dioxide (SO_2), hydrogen sulfide (H_2S), carbon monoxide (CO), hydrogen chloride (HCl) and so on [Sigurdsson et
 640 al., 2015]. We think there are enough studies to prove that these gases which leak from underground fissure at pre-
 641 eruption period can change the temperature of upper atmosphere and lower the electron density [Cnossen, 2022;
 642 Qian et al., 2013; Roble & Dickinson, 1989]. In the meantime, the volcanic eruption phase will generate much more
 643 gravity wave, seismic wave, acoustic wave and infrasonic wave [Ricardo Garza-Girón, 2023]. Although this effect

644 is not global and has low strength, the local variation will propagate outward by the dynamic action of the upper
 645 atmosphere.

646



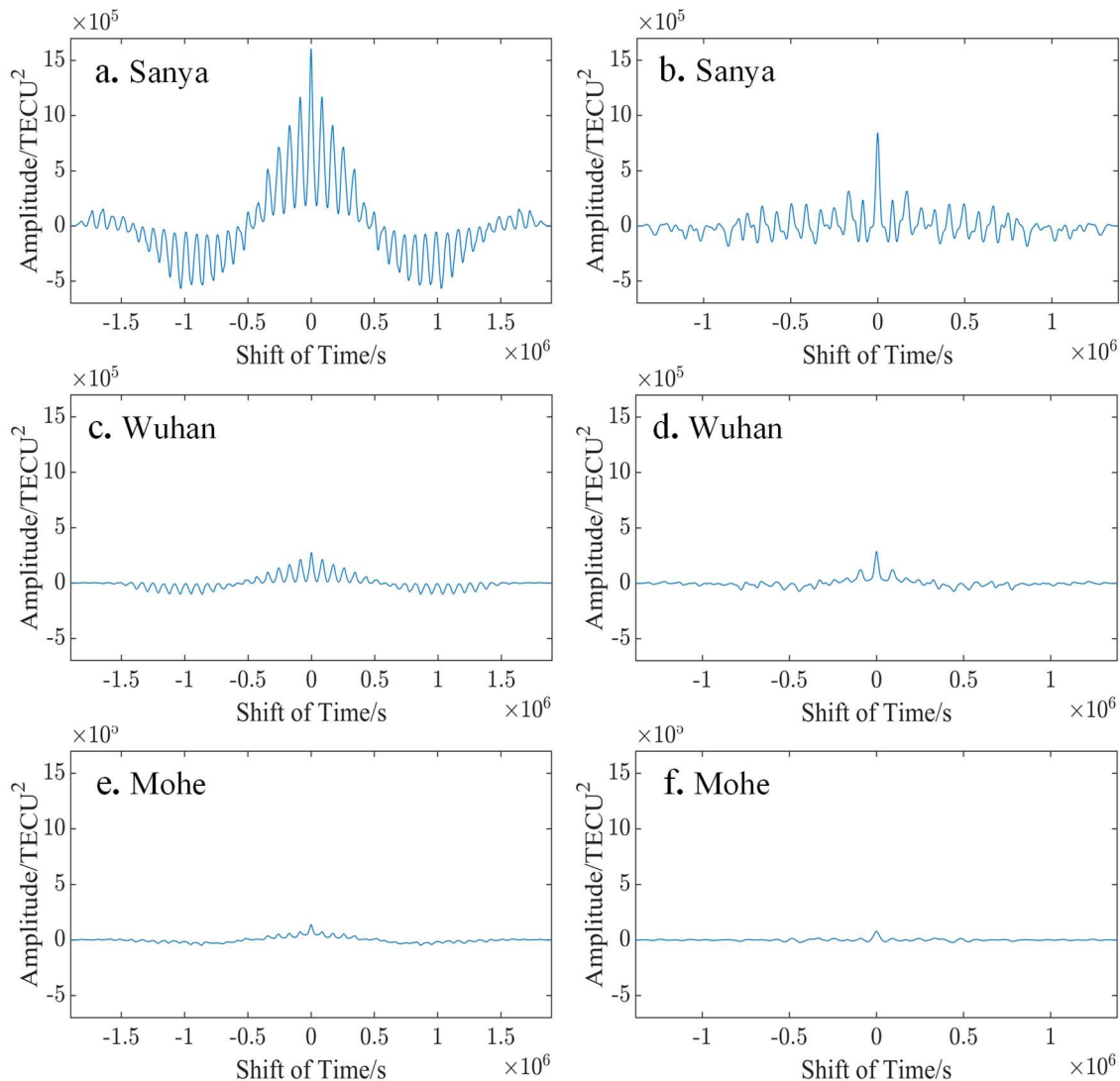
647 **Figure 2.** Plot of ionospheric TEC values over Sanya station, Hainan, during the volcanic eruption. The green
 648 dashed lines point out the median values for the time period, the blue dashed lines indicate the maximum and minor
 649 values, and the red solid lines mark the fitted value over the station. (a–l): TEC variation during December 10–21,
 650 2021 (Doy344–Doy355). It can be found that on December 20th, at the beginning of Doy354, the fitted value
 651 deviates significantly from the median value, and the TEC fluctuates distinctly. The maximum peak reaches 50
 652 TECU, which is higher than the daily maximum peak of about 20–30 TECU. (m–t): The TEC variation from 10 to
 653 17 January 2022 (Doy10–Doy17). It can be found that there are fluctuations in ionosphere 15 TEC, but the overall
 654 anomaly is smaller than that in (a–l).



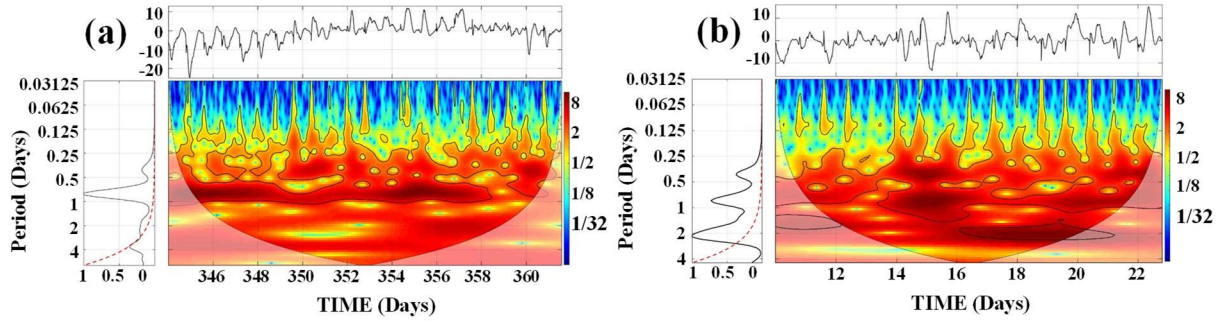
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656 **Figure 3.** The TEC anomalies over Sanya, Wuhan and Mohe stations during the eruption phase. The red solid line is
 657 the interpolation of the fitting value, and the blue horizontal line indicates the zero level of TECU. The blue
 658 asterisk is the maximal value of the similar waveform selected. (a): Anomalies during the period from December
 659 10 to 31, 2021 (Doy344–Doy365) over Sanya station. (b): Plot of anomalous variation from January 10 to 17, 2022
 660 (Doy10–Doy17) over Sanya station. (c): Anomalies during the period from December 10 to 31, 2021 (Doy344–
 661 Doy365) over Wuhan Zuoling station. (d): Plot of anomalous variation from January 10 to 17, 2022 (Doy10–Doy17)
 662 over Wuhan Zuoling station. (e): Anomalies during the period from December 10 to 31, 2021 (Doy344–Doy365)
 663 over Mohe station. (f): Plot of anomalous variation from January 10 to 17, 2022 (Doy10–Doy17) over Mohe station.
 664 The TEC decays substantially from 10 days before the 20th (Doy354), and the anomalous variations exhibit
 665 significant increase with oscillations before the outbreak on the 20th. Similar decay can be found in several days
 666 before the outbreak on 13th, but the overall amplitude is smaller than that of (a, c, and e). After decaying, TEC
 667 anomalies appear to undergo significant enhancement on the Doy354 and Doy15, respectively. Here is the
 668 difference: there are many peaks with different shapes and intensities during the intermittent stage. However, the
 669 anomalies variation in main stage has similar distinct waveform characteristic and relatively close intensity.

670



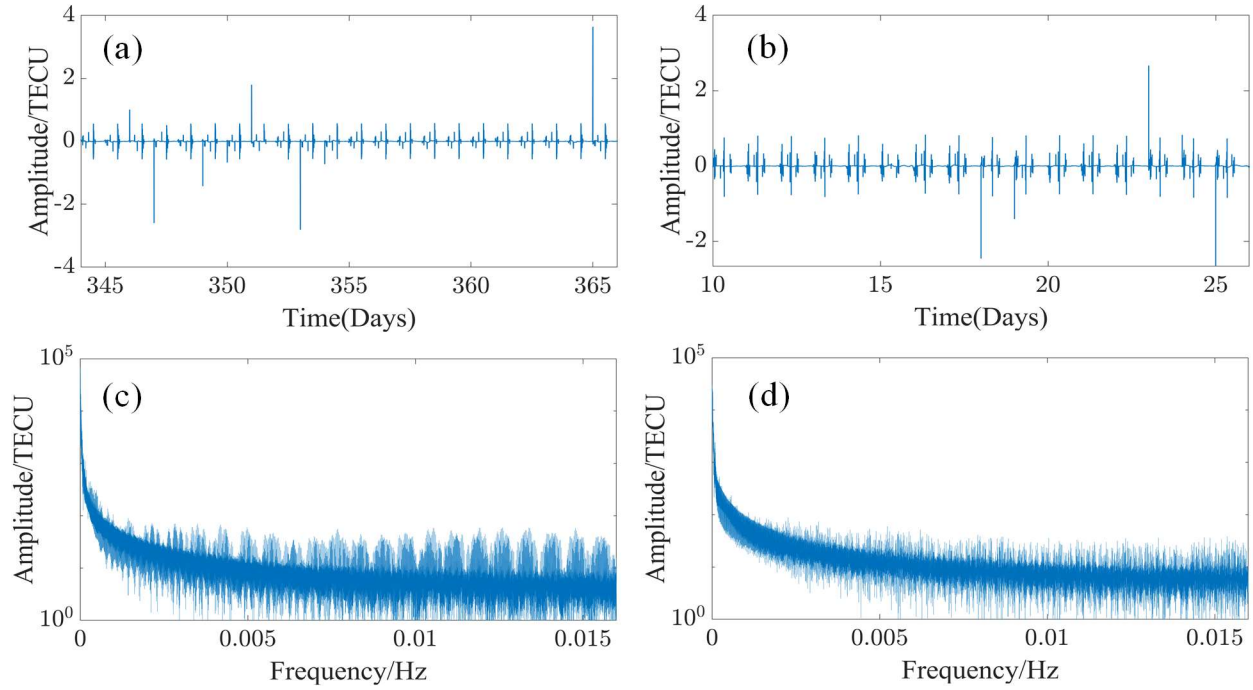
671
 672 **Figure 4.** Autocorrelation results of anomalous variations during the eruption. **The horizontal coordinates are the**
 673 **time-shift taken by the autocorrelation function.** (a): The autocorrelation of the intermittent stage over Sanya shows
 674 that the maximum peak value of the autocorrelation function is not prominent. (b): It can be found that the
 675 maximum peak of autocorrelation function is 4–10 times higher than other smaller peaks during the main stage over
 676 Sanya. (c): The result of autocorrelation analysis of the TEC anomaly during the intermittent stage over Wuhan
 677 Zuoling station. Its signal attenuation degree is between that over Sanya and Mohe, with similar change pattern. (d):
 678 The plot of autocorrelation analysis over Wuhan Zuoling station during the main stage. (e): The plot of
 679 autocorrelation analysis of the intermittent stage over Mohe station. The the attenuation of the signal intensity is
 680 shown obviously, with similar variation characteristics as Sanya. (f): The result of autocorrelation of the main stage
 681 over Mohe. **By demonstrating a clear ionospheric response to eruptions in this study, we can reduce the noise and**
 682 **increase the peak value through autocorrelation. After autocorrelating, each peak of the result may show a peak of**
 683 **TEC anomalies for one eruption. We can get an approximate number of eruptions by this method. Our work may be**
 684 **able to provide a kind of corroboration for volcanic research in such data-missing case.**



685

686 **Figure 5.** The plot through the Wavelet analysis of the anomalies. (a): The result of wavelet analysis of the TEC
 687 anomaly over Sanya during the intermittent stage. (b): The result of wavelet analysis of the TEC anomaly over
 688 Sanya during the main stage. For (a) and (b), the upper panels show the results of the low-frequent component of the
 689 signals. The lower panels exhibit the continuous wavelet transform time-frequency diagram without the edge effect.
 690 The left panels reveal the signal period extracted through wavelet transform.

691



692

693 **Figure 6.** The supplementary plots based on the Fourier analysis of the anomalies. (a): High-frequency part of the
 694 discrete wavelet transform of the TEC abnormal signals over Sanya during the intermittent phase. (b): High-
 695 frequencies part over Sanya during the main outburst phase. (c): During December 10–21, 2021 (Doy344–Doy355),
 696 the energy is mainly concentrated in the frequency range of an order of $0 - 10^{-3}$ Hz over Sanya. (d): The
 697 amplitude spectrum from 10 to 17 January 2022 (Doy10–Doy17) over Sanya. It can be discerned that the energy is
 698 mainly concentrated in $0 - 10^{-3}$ Hz.

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700